

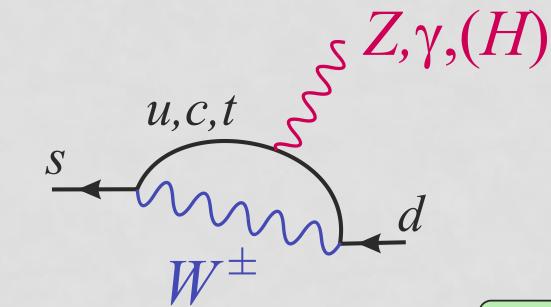
# $K_L \rightarrow \pi \ell\ell$ and radiative $K$ decays

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in collaboration with  
Christopher Smith (IPNL)



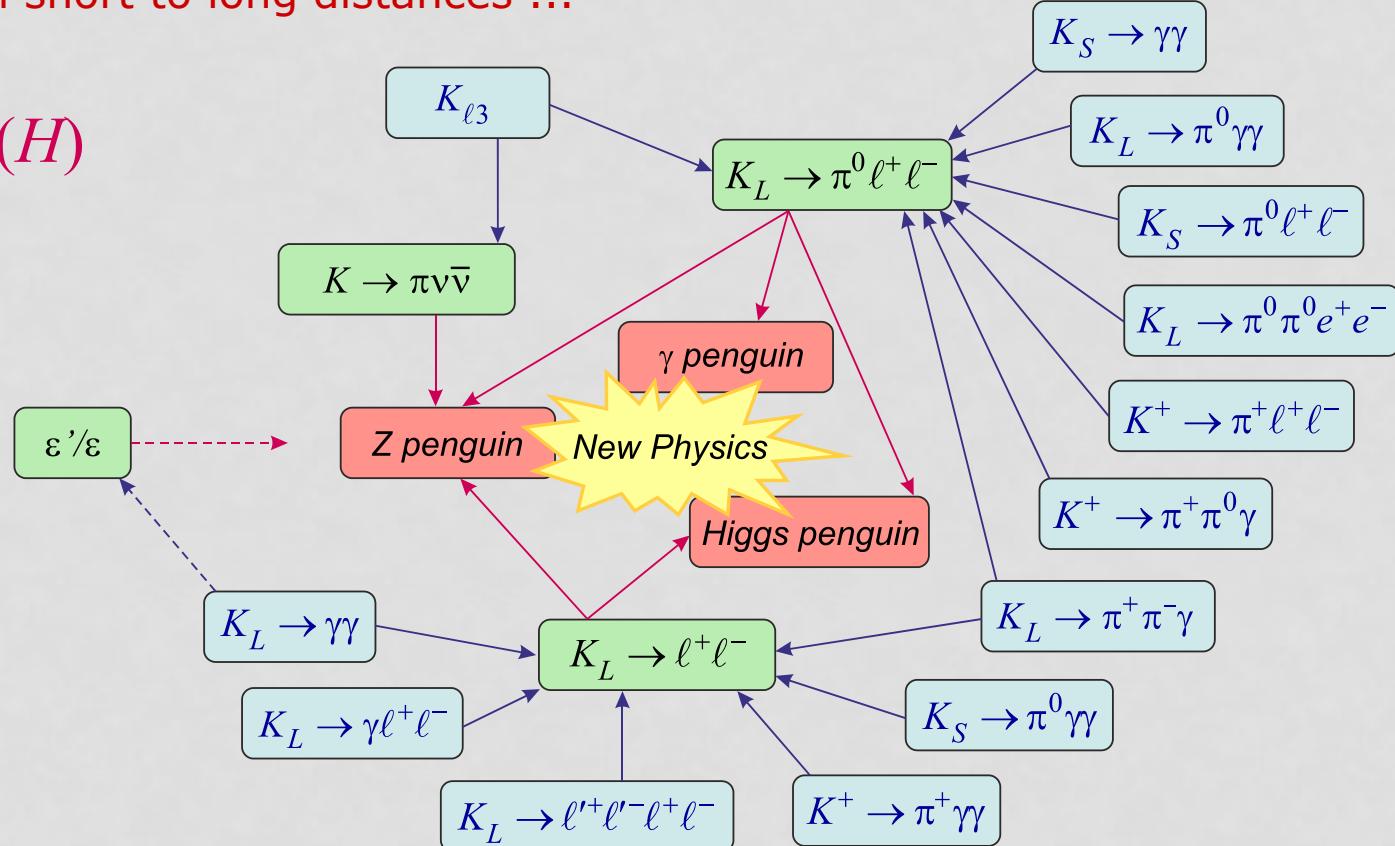
2012 Project X Physics Study, Kaon Working Group  
Fermi National Accelerator Laboratory

## A. Round-trip from short to long distances ...

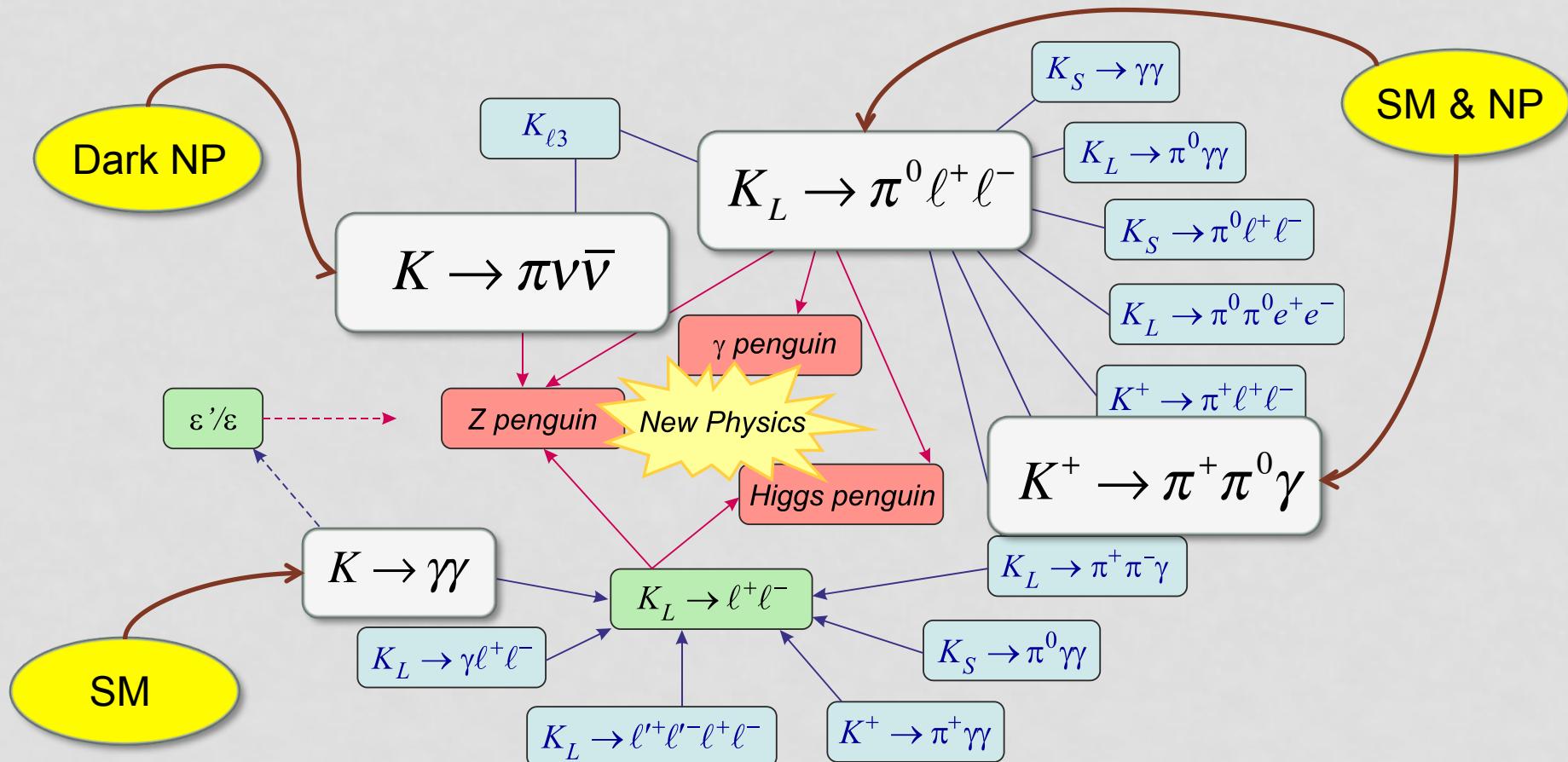


$\approx 1000 \text{ GeV}$   
New Physics

$\approx 100 \text{ GeV}$   
Electroweak windows



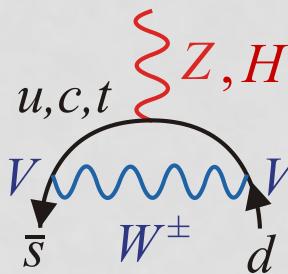
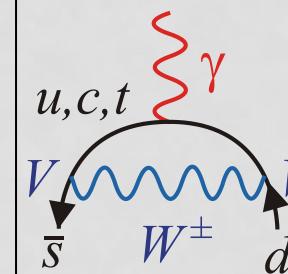
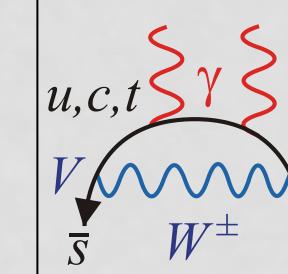
## B. The « connections and complementarity » picture :



TH : Ulrich's, Joachim's, Stephanie's and Wolfgang's talk.

EXP : Patrizia's, David's, Elizabeth's, Yau's and Laurence's talk.

## C. Disentangling FCNC :

			
Rare	$K \rightarrow \pi \nu \bar{\nu}$	SD (no H)	-
	$K_L \rightarrow \pi^0 \ell^+ \ell^-$	SD	$\mathcal{E}_K$ LD
	$K_{L,S} \rightarrow \ell^+ \ell^-$	SD	$\alpha_{QED}$ LD
Radiative	$K_S, K^+ \rightarrow \pi^{0,+} \ell^+ \ell^-$	Negligible, except for CP- asymmetries.	LD  ( $\alpha_{QED}$ -suppr.)
	$K \rightarrow (n\pi)\gamma\gamma$	-	LD
	$K \rightarrow (n\pi)\gamma$	LD	-

The dark side of  $K \rightarrow \pi\nu\bar{\nu}$

## A. Are there only SM particles at low-energy?

Experimentally:

- Even very light states could be missed if **very weakly interacting**,
- There is **dark matter** in the Universe; it could be relatively light.

Theoretically: Plenty of models predict new light particles.

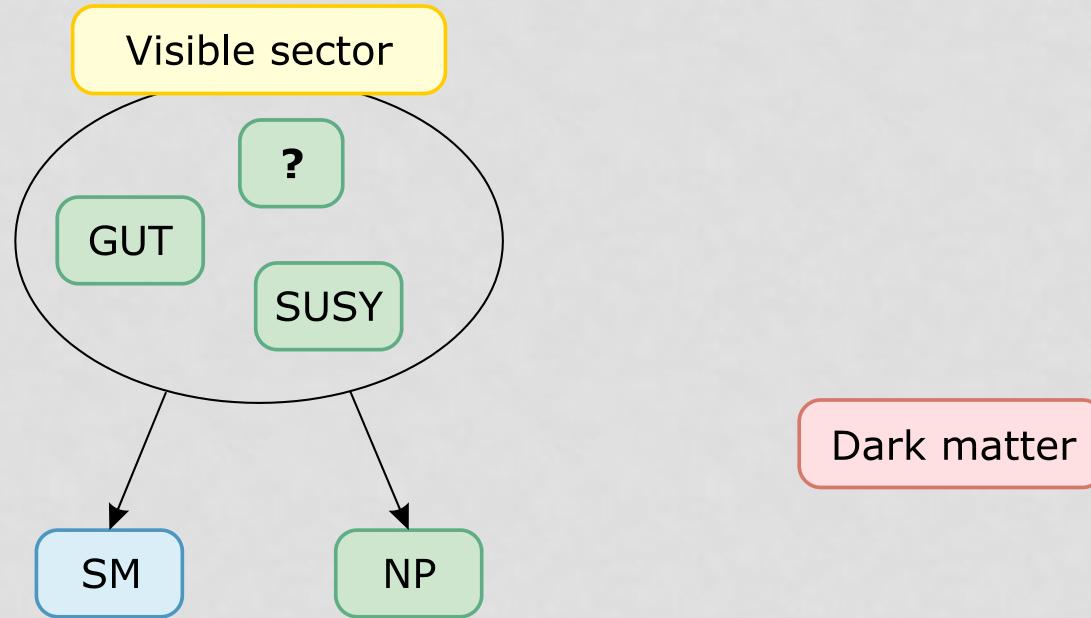
- **Pseudo-Goldstone scalars** (axion, familon,...),
- **U(1) vectors** (string, ED,...),
- **Hidden sectors & messengers** (SUSY, mirror worlds,...)
- Many others: **millicharged fermions, dilaton, majoron, neutralino, sterile neutrino, gravitino,...**

Phenomenologically, the dark state can be assumed

- **Unstable** → No DM constraints,
- **Long-lived** → Escape as missing energy,
- **Weakly coupled** → SM processes unaffected.

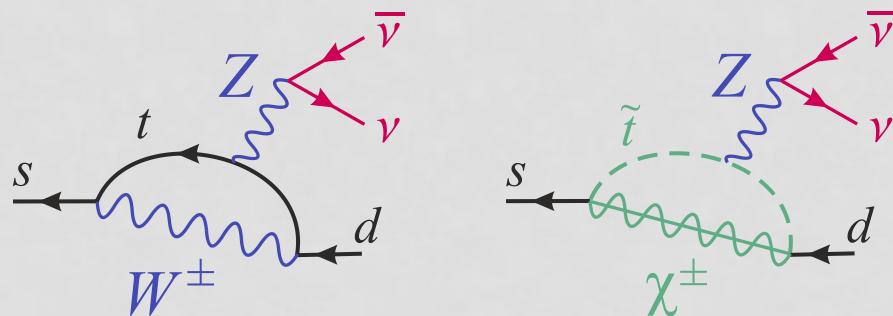
Main impact is to  
open new decay  
channels.

## B. How to systematically investigate the low-energy particle content?

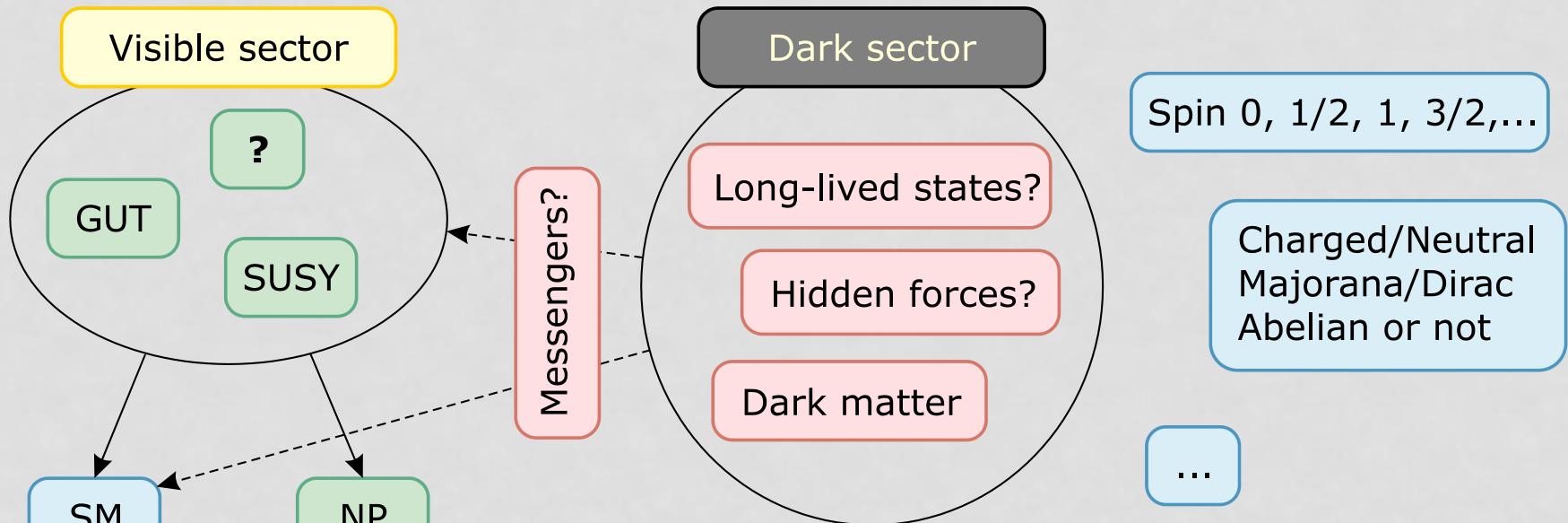


$$\mathcal{L}_{SM} + \frac{c_i}{\Lambda^2} Q_i + \dots$$

Buchmüller, Wyler '86  
Weinberg '79



## B. How to systematically investigate the low-energy particle content?

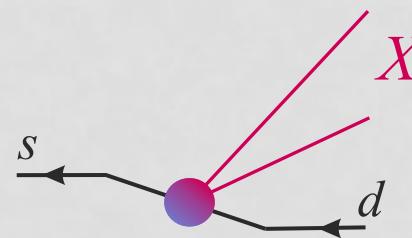
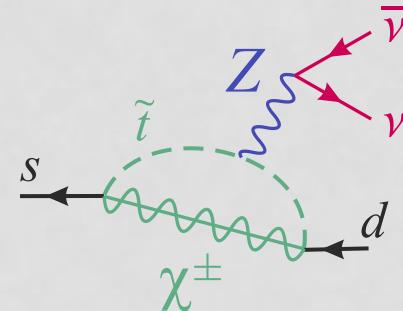
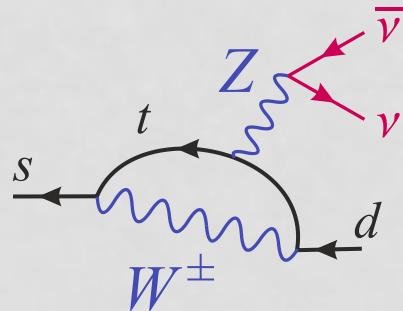


$$\mathcal{L}_{SM}$$

$$+ \frac{c_i}{\Lambda^2} Q_i + \dots$$

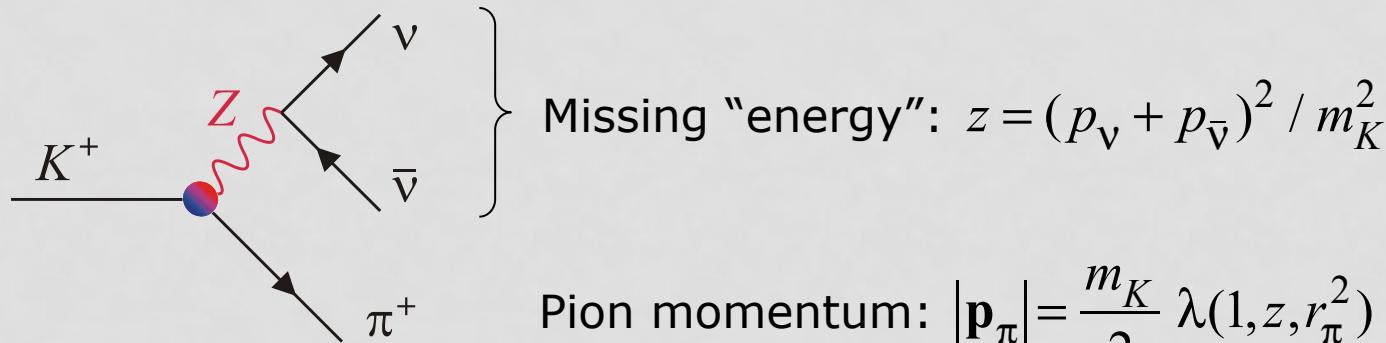
$$+ \sum_{d \geq 3} \frac{c_i}{\tilde{\Lambda}^{d-4}} Q'_i + \dots$$

Kamenik, Smith '11



### C. What rare K decays could tell?

Only the pion is seen, whose energy is not fixed (three-body decay).

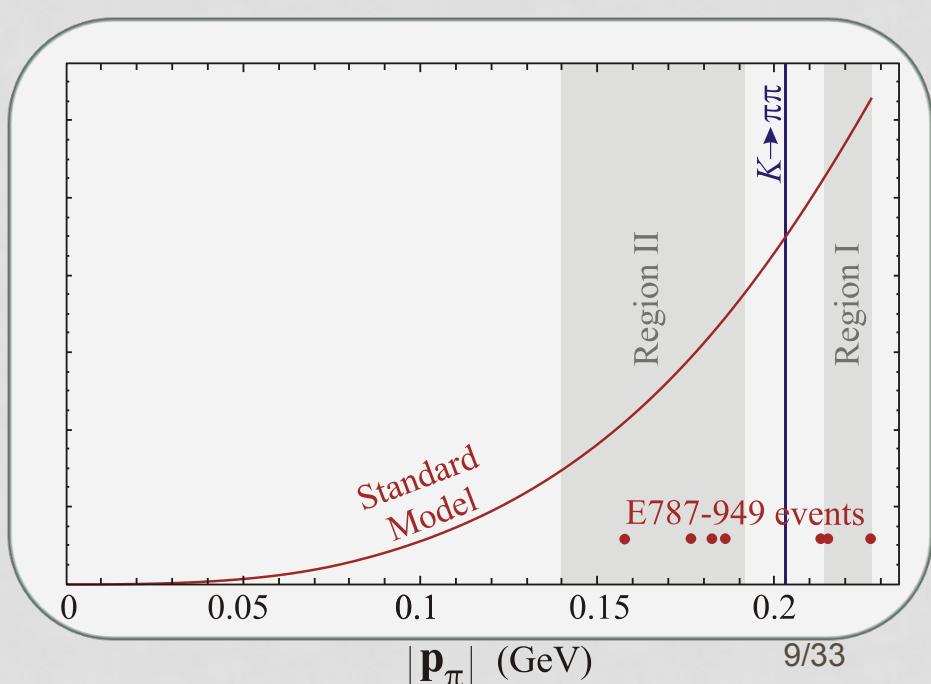


The V-A interaction of the SM predicts:

$$\frac{\partial \ln \Gamma}{dz} \sim \frac{|\mathbf{p}_\pi|^3}{m_K^3} |f(z)|^2$$

with  $\langle \pi | \bar{s} \gamma^\mu d | K \rangle \approx f(z)(p_K + p_\pi)^\mu$ .

Crucial tool for background rejection!



### C. What rare K decays could tell? Five messages:

1- For  $K \rightarrow \pi v\bar{v}$ , use the **correct SM spectrum**:

Mescia & Smith '06

- Fixed from  $K_{\ell 3}$  slopes  $\rightarrow f(z)$  will evolve!

2- For  $K \rightarrow \pi + E$  with  $E = \text{dark state(s)}$ , the **spectrum may not be V-A**.

- It could even be a two-body decay (e.g. axion-like scalars).
- So, keep open the possibility to alter the spectrum,
- Measure  $\Gamma_{\text{Region I}} / \Gamma_{\text{Region II}}$  as model-independently as possible.

3- Dark particles may be **more massive than neutrinos**.

- Important to probe the region between  $\pi\pi\pi$  and the  $\pi\pi$  peak.

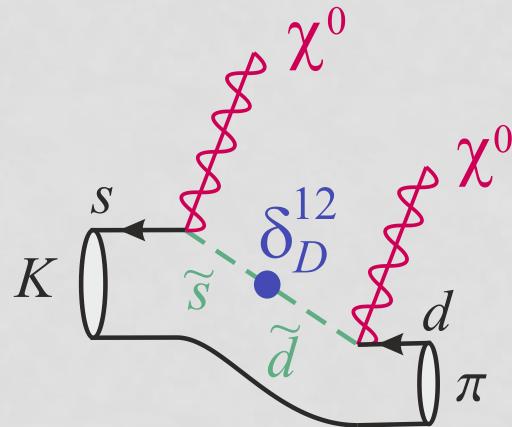
4-  $K_L \rightarrow \gamma + E$  is also **very sensitive** to dark states.

- Neutrino mode not a background since  $B(K_L \rightarrow \gamma v\bar{v})^{SM} \approx 3 \cdot 10^{-13}$ .

5-  $K_L \rightarrow E$  is **best** but inaccessible (?);  $K \rightarrow \pi\pi + E$  are not competitive.

## Example 1: Very light neutralinos in the generic MSSM

Dreiner et al. '09  
Kamenik, Smith '11

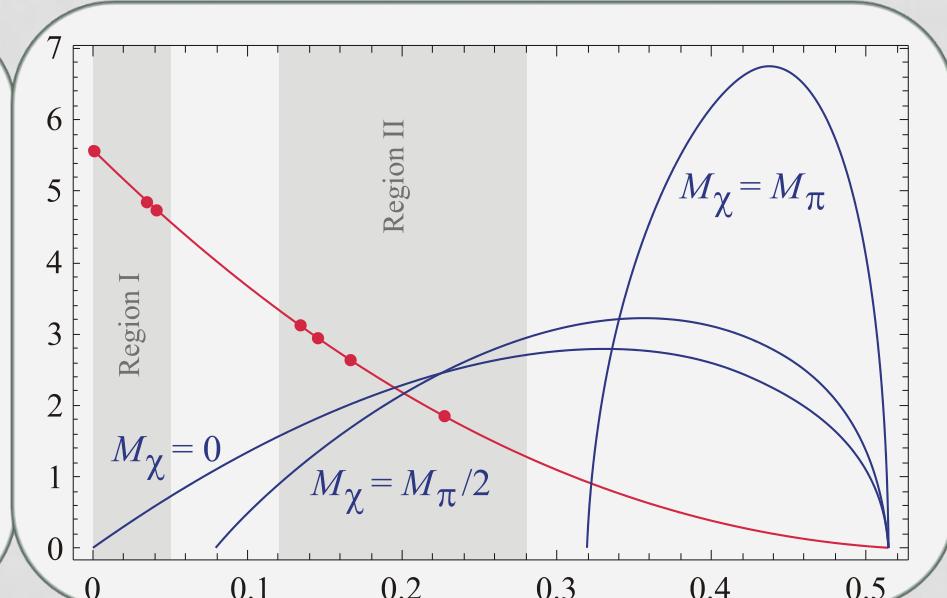
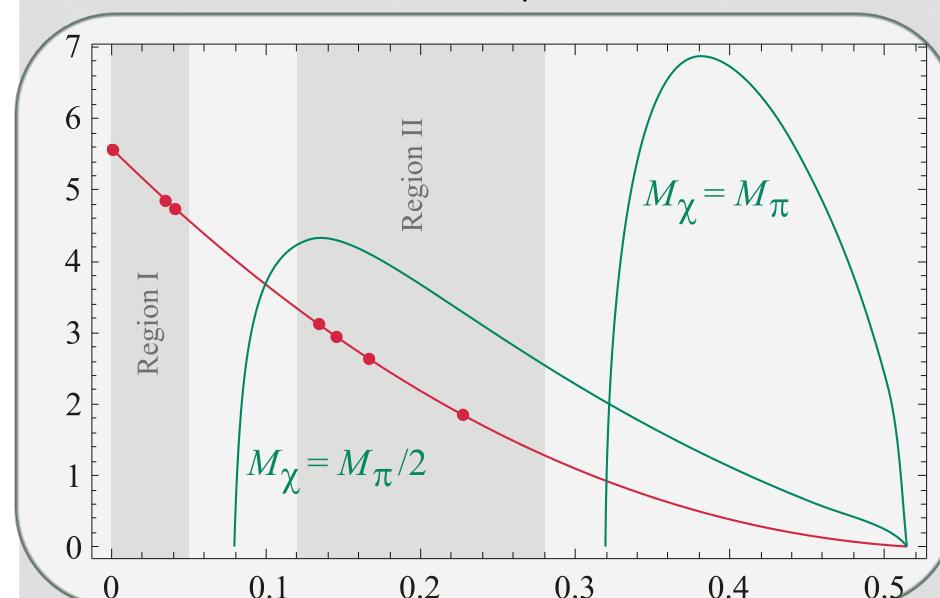


(Axial-)vector :

$$\bar{s} \gamma^\mu (1 \pm \gamma_5) d \otimes \bar{\chi} \gamma_\mu \gamma_5 \chi$$

(Pseudo)scalar:

$$\bar{s} (1 \pm \gamma_5) d \otimes \bar{\chi} (1 \pm \gamma_5) \chi$$



## Example 2: Kinetically-mixed light vector boson

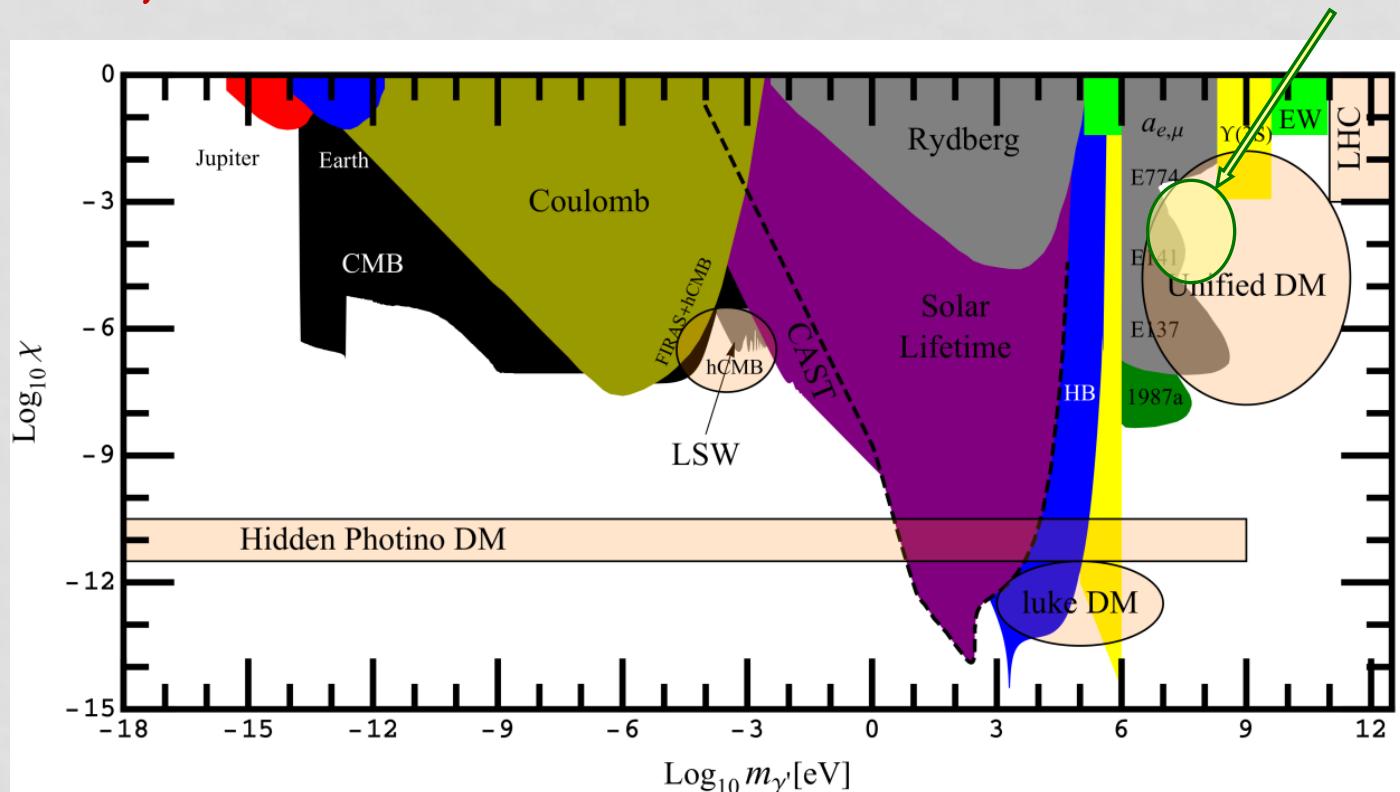
Holdom '86  
Arkani-Hamed et al. '08  
Kamenik, Smith '11

$$\mathcal{L}_{kin} = \frac{\chi}{2} B_{\mu\nu} \times V^{\mu\nu} \rightarrow e\chi V_\mu \left( \frac{2}{3}\bar{u}\gamma^\mu u - \frac{1}{3}\bar{d}\gamma^\mu d - \frac{1}{3}\bar{s}\gamma^\mu s \right)$$

Up to  $m_V \neq 0$ , rates proportional to those of radiative decays.

$K \rightarrow \pi V$  : Very suppressed because  $K \rightarrow \pi\gamma$  is forbidden.

$K \rightarrow \gamma V$  : A bound in the  $10^{-12}$  range probes down to  $\chi$  :  $4 \cdot 10^{-5}$ .

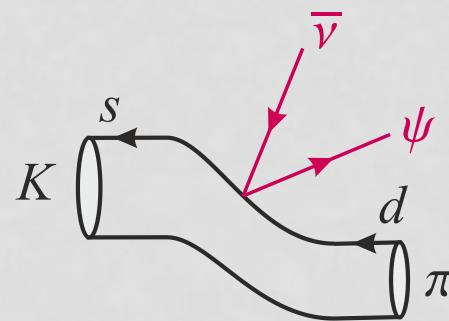


Jaeckel, Ringwald, 1002.0329

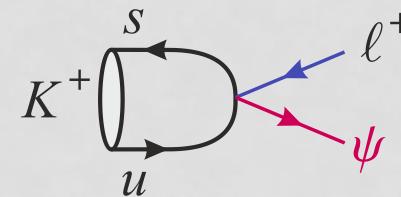
### Example 3: Leptonic dark state (sterile neutrino,...)

Correlated neutral and charged current interactions:

$$\mathcal{H}_{\text{eff}} = \frac{1}{\Lambda^2} (\bar{d}_L s_R \otimes \bar{\nu}_L \psi + \bar{u}_L s_R \otimes \bar{\ell}_L \psi)$$



Rare FCNC modes  
like  $K \rightarrow \pi + E$  probe  
scales up to  $\Lambda$ : 100 TeV.



Current universality test,

$$\mathcal{R}_K^{\text{exp}} = \frac{\Gamma(K_{e2})^{\text{SM}} + \Gamma(K \rightarrow e\psi)}{\Gamma(K_{\mu 2})^{\text{SM}} + \Gamma(K \rightarrow \mu\psi)}$$

reaches a similar scale  $\Lambda$ : 80 TeV .

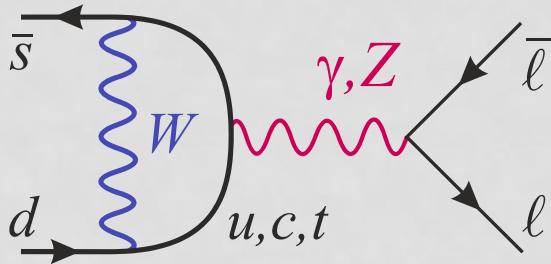
( $\delta R_\pi^{\text{exp}}$  :  $10^{-3}$  would probe  $\Lambda$ : 70 TeV)

$$K_L \rightarrow \pi^0 \ell^+ \ell^-$$

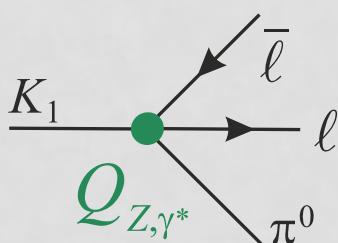
$$K \rightarrow \pi \ell^+ \ell^- 1/6$$

$$K_L \sim K_2 + \varepsilon K_1 \rightarrow \pi^0 \ell^+ \ell^-$$

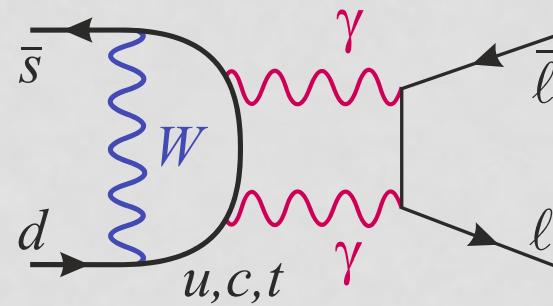
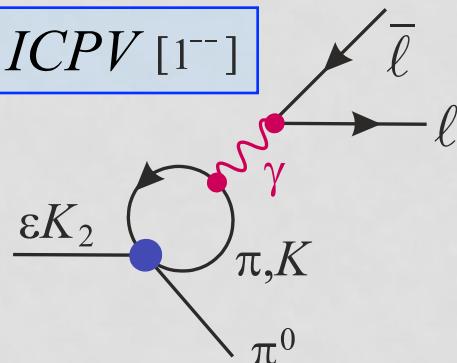
A. Various contributions :



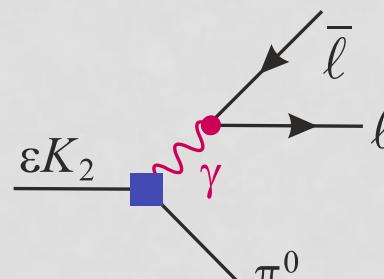
**DCPV**  $[0^{-+}, 1^{\pm\pm}]$



**ICPV**  $[1^{--}]$



**CPC**  $[0^{++}, 2^{++}]$



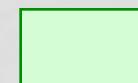
Direct CPV

(Short-distance)



Indirect CPV

(Long-distance)



CPC

(Long-distance)

## B. Direct CPV: Matrix elements of the dimension-six operators

Mescia, Smith '06

LD effects for the top/charm “pure” SD contribution = matrix elements

$$Q_{eff}^V = (\bar{s}d)_V \otimes (\bar{\ell}\ell)_V, \quad Q_{eff}^A = (\bar{s}d)_V \otimes (\bar{\ell}\ell)_A$$

As for  $K \rightarrow \pi v\bar{v}$ , those are extracted from  $K_{\ell 3}$  decays:

		$t_+$	$f(0)$	slopes	$r_K$	$r$	Future?
$\kappa_e^{V,A}$	0.7691(64)	-	77%	12%	9%	2%	$\pm 0.0046$
$\kappa_\mu^V$	0.1805(16)	-	73%	16%	8%	2%	$\pm 0.0011$
$\kappa_\mu^A$	0.4132(51)	-	54%	38%	6%	2%	$\pm 0.0031$

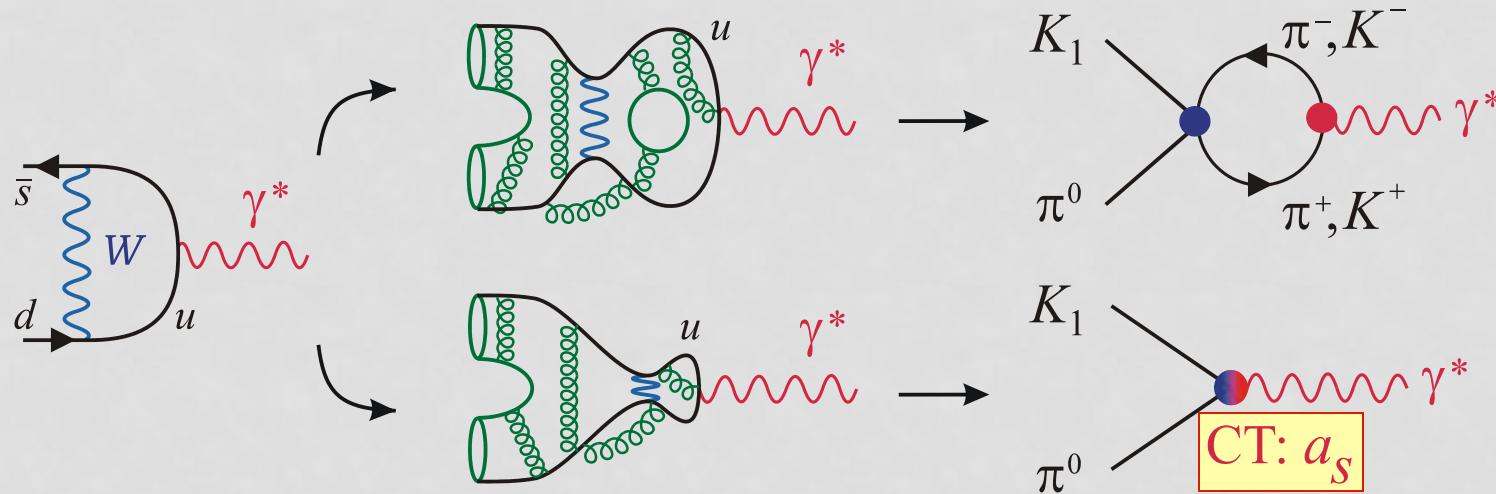
$$\kappa_\ell^{V,A} \sim \int d\Phi_3 |\langle \pi^0 \ell \bar{\ell} | Q_{eff}^{V,A} | K_L \rangle|^2$$

Already very precise compared the other contributions.

### C. Indirect CPV: Long-distance photon penguin

D'Ambrosio et al. '98

Indirect CP-violation is  $K_L \rightarrow \varepsilon K_1 \rightarrow \pi^0 \ell^+ \ell^-$ , related to  $K_S \rightarrow K_1 \rightarrow \pi^0 \ell^+ \ell^-$ :



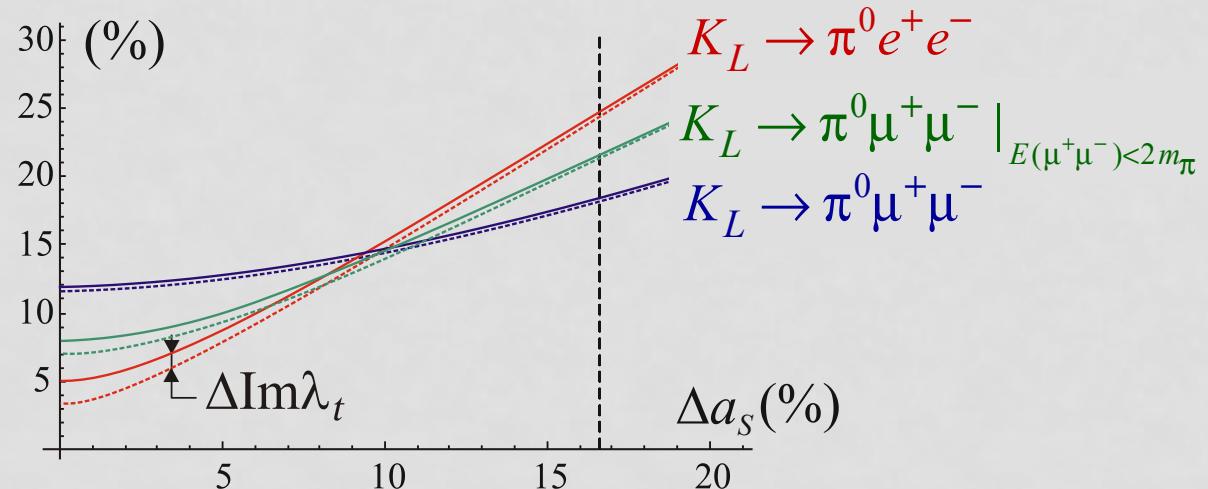
Loops are rather small, a single counterterm  $a_S$  dominates.

It is fixed from  $K_S \rightarrow \pi^0 \ell^+ \ell^-$  (up to its sign) measured by NA48:

$$\left. \begin{aligned} Br(K_S \rightarrow \pi^0 e^+ e^-)_{m_{ee} > 165 \text{ MeV}} &= (3.0^{+1.5}_{-1.2} \pm 0.2) \times 10^{-9} \\ Br(K_S \rightarrow \pi^0 \mu^+ \mu^-) &= (2.9^{+1.4}_{-1.2} \pm 0.2) \times 10^{-9} \end{aligned} \right\} \rightarrow |a_S| = 1.2 \pm 0.2$$

### C. Indirect CPV: Long-distance photon penguin

This CT is the main source of error for



Besides  $K_S \rightarrow \pi^0 \ell^+ \ell^-$ , the paths to constrain or measure  $a_S$  are (among other):

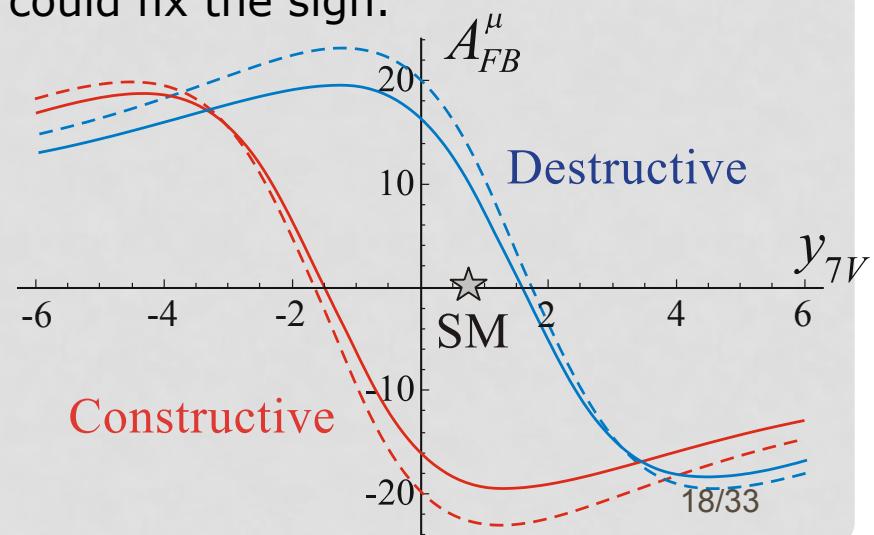
- FB asymmetries for  $K_L \rightarrow \pi^0 \mu^+ \mu^-$  could fix the sign.

$$A_{FB}^\ell = \frac{N(E_- > E_+) - N(E_- < E_+)}{N(E_- > E_+) + N(E_- < E_+)}$$

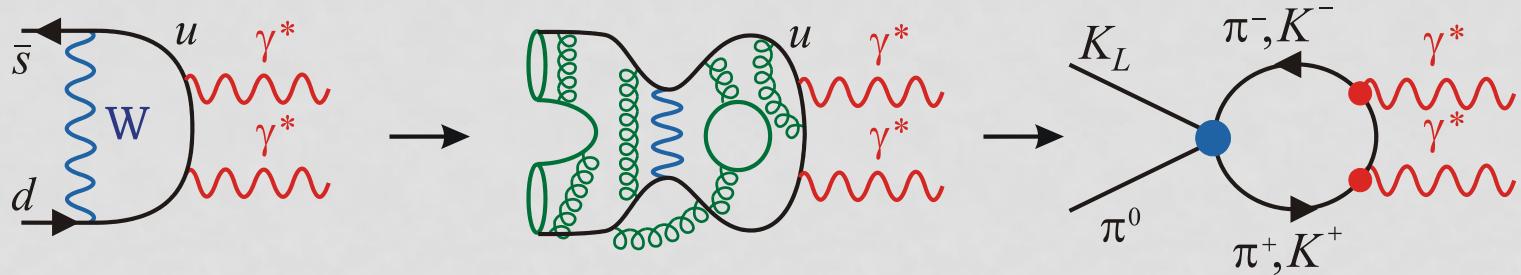
Clean enough to fix the sign of  $a_S$

Mescia,Trine, Smith '06

$$(y_{\gamma V} Q_{eff}^V = y_{\gamma V} (\bar{s}d)_V \otimes (\bar{\ell}\ell)_V)$$

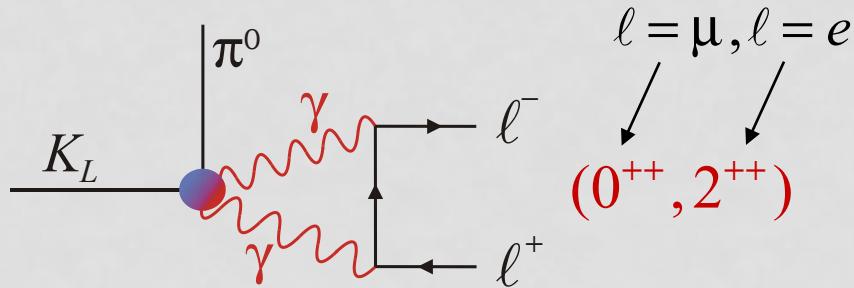


## D. CPC: Long-distance double photon penguin



LO ( $p^4$ ) is finite, produces  $\ell^+ \ell^-$  in a scalar state only (helicity-suppressed),

Higher order estimated using the  $K_L \rightarrow \pi^0 \gamma \gamma$  rate and spectrum:



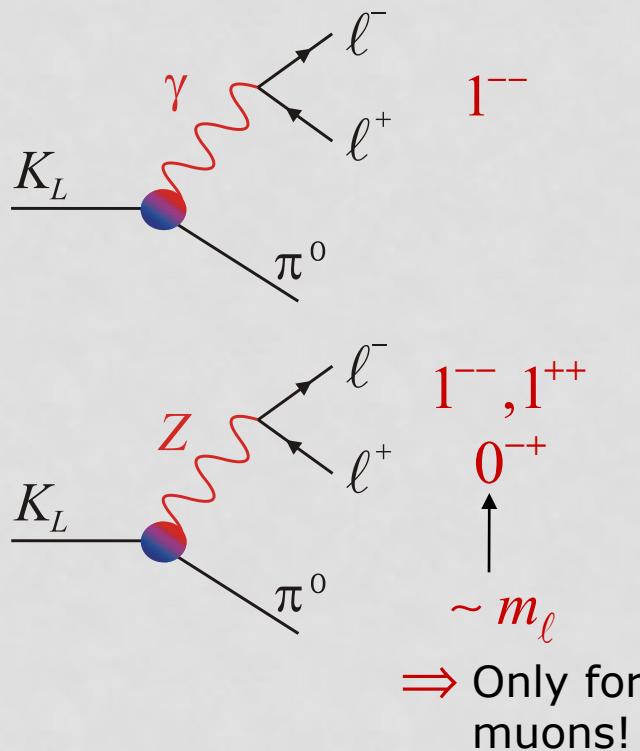
- Production of  $(\mu^+ \mu^-)_{0^{++}}$  under control within 30%. Isidori,Unterdorfer,Smith '04
- No signal of  $(\gamma \gamma)_{2^{++}}$  implies  $(e^+ e^-)_{2^{++}}$  is negligible. Buchalla,D'Ambrosio, Isidori '03  
( $K_S \rightarrow \gamma \gamma$  is also useful to constrain the  $p^6$  CT structure)

*In the SM :*

	$V, A$	$K^0 - \bar{K}^0$	$2^{++}$	$0^{++}$	SM ( $\times 10^{-11}$ )	Experiment
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	<b>100%</b>	( $\approx 1\%$ )	-	-	$2.43_{-0.40}^{+0.40}$	$< 2.6 \cdot 10^{-8}$ E391a
$K_L \rightarrow \pi^0 e^+ e^-$	<b>40%</b>	60%	(<3%)	-	$3.23_{-0.79}^{+0.91}$	$< 2.8 \cdot 10^{-10}$ KTeV
$K_L \rightarrow \pi^0 \mu^+ \mu^-$	<b>30%</b>	35%	-	35%	$1.29_{-0.23}^{+0.24}$	$< 3.8 \cdot 10^{-10}$ KTeV
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	<b>100%</b>	-	-	-	$7.81_{-0.80}^{+0.80}$	$1.73_{-1.05}^{+1.15} \cdot 10^{-11}$ E787 E949

And what's beyond ?

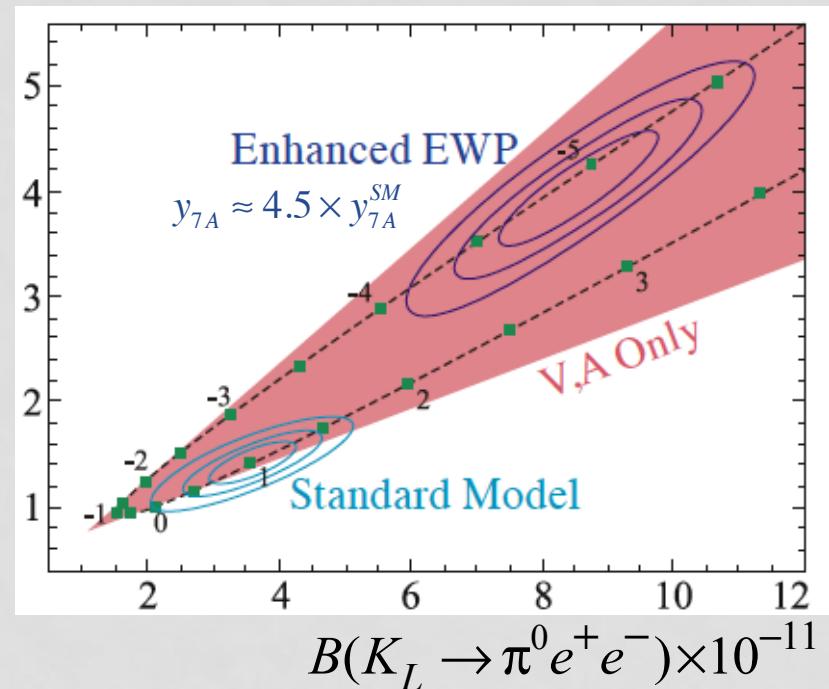
The  $K_L \rightarrow \pi^0 \ell^+ \ell^-$  modes permit to isolate the  $\gamma$  penguin:



Specific regions in the plane for specific correlations between  $\gamma$  and  $Z$  penguins.

Isidori,Unterdorfer,Smith '04

$$B(K_L \rightarrow \pi^0 \mu^+ \mu^-) \times 10^{-11}$$

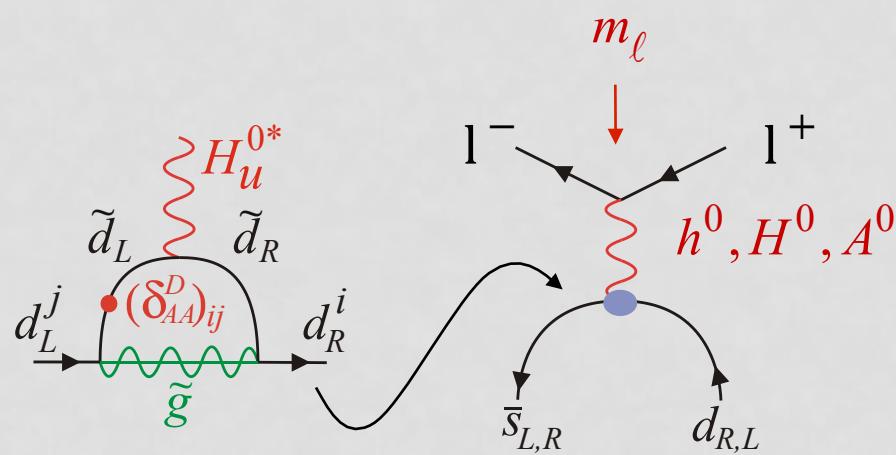


$$\Leftrightarrow 0.1 + 0.24 B_e \leq B_\mu \leq 0.6 + 0.58 B_e$$

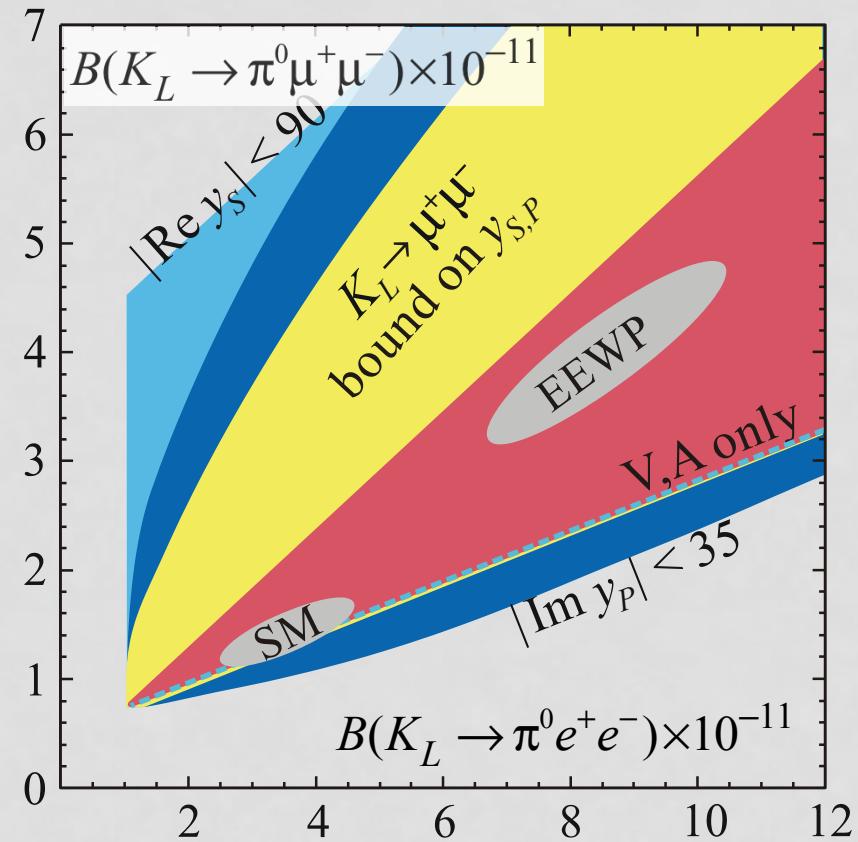
$$\text{with } B_\ell \equiv B(K_L \rightarrow \pi^0 \ell^+ \ell^-) \cdot 10^{11}$$

They also probe a broad class of helicity-suppressed NP effects, as those generated by **neutral Higgs penguins**.

Isidori, Retico '01, '02  
Mescia, Trine, Smith '06



$$H_{eff} \sim y_S (\bar{s}d)(\bar{\ell}\ell) + y_P (\bar{s}d)(\bar{\ell}\gamma_5\ell)$$



Even if directly correlated with  $K_L \rightarrow \mu^+ \mu^-$ , large effects still possible.

Looking for the unexpected?

1-  $K_L \rightarrow \pi^0 \ell^+ \ell^-$  probe the (pseudo-)scalar and (pseudo-)tensor operators:

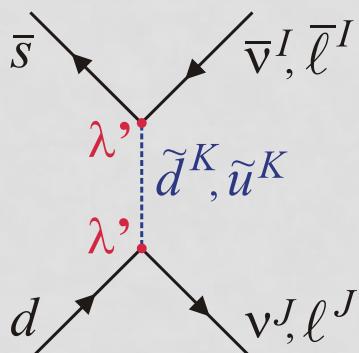
$$H_{eff} \sim y_S^{IJ} (\bar{s}d)(\bar{\ell}^I \ell^J) + y_P^{IJ} (\bar{s}d)(\bar{\ell}^I \gamma_5 \ell^J) + y_T (\bar{s}\sigma_{\mu\nu}d)(\bar{\ell}\sigma^{\mu\nu}\ell) + \dots$$

Presumably killed by  $K_L \rightarrow ee, e\mu$       ...but tensors still free.

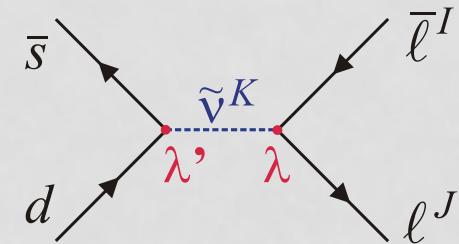
2- Similar operators for  $K \rightarrow \pi \nu \bar{\nu}$  require active right-handed neutrinos.

3- The RPV couplings can induce quark & lepton flavor transitions,  
hence could contribute to all  $P \rightarrow P' \nu^I \bar{\nu}^J$  and  $P \rightarrow P' \ell^I \bar{\ell}^J$  decays :

$$W_{RPV} = \underbrace{\mu' L^I H_d}_{\Delta L = 1} + \underbrace{\lambda^{IJK} L^I L^J E^K + \lambda'^{IJK} L^I Q^J D^K + \lambda''^{IJK} U^I D^J D^K}_{\Delta B = 1}$$



$$\bar{s} \gamma^\mu (1 \pm \gamma_5) d \otimes \begin{cases} \bar{\nu} \gamma_\mu (1 - \gamma_5) \nu \\ \bar{\ell} \gamma_\mu (1 - \gamma_5) \ell \end{cases}$$



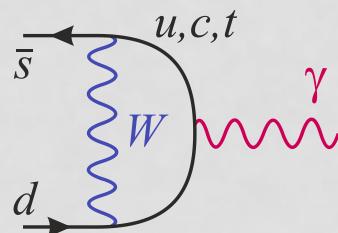
$$\bar{s} (1 \pm \gamma_5) d \otimes \bar{\ell} (1 \mp \gamma_5) \ell$$

$$s \rightarrow d\gamma$$

## A. The flavor-changing electromagnetic operators

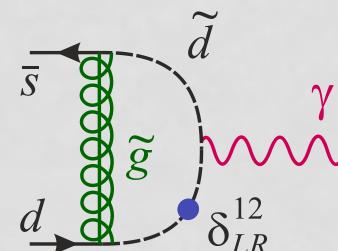
$$\begin{aligned}
 \mathcal{H}_{eff} = & C_{\gamma}^{L,R} \bar{s}_{R,L} \sigma^{\mu\nu} d_{L,R} F_{\mu\nu} & \text{Dim 5} \\
 & + C_{\gamma^*}^{L,R} \bar{s}_{L,R} \gamma^\nu d_{L,R} \partial^\mu F_{\mu\nu} & \text{Dim 6} \\
 & + C_{\gamma\gamma}^{L,R} \bar{s}_{R,L} d_{L,R} F_{\mu\nu} F^{\mu\nu} + \tilde{C}_{\gamma\gamma}^{L,R} \bar{s}_{R,L} d_{L,R} F_{\mu\nu} \tilde{F}^{\mu\nu} & \text{Dim 7} \\
 & + \dots
 \end{aligned}$$

Induced at one loop in the SM ( $x_q \equiv m_q^2 / M_W^2$ ):



Magnetic:  $C_{\gamma}^{L,R} : m_{s,d} \sum_{q=u,c,t} V_{qs}^* V_{qd} D'(x_q)$

Sensitive to NP: Dimension 5 & alternative chirality flips



## A. Best windows for $s \rightarrow d\gamma$

The magnetic operators  $Q_\gamma^{L\pm R}$  contribute to all the radiative modes:

$$K \rightarrow (n\pi)(m\gamma^{(*)}), \quad \begin{array}{ccc} n = & 0 & 1 & 2,3 \\ m = & 2,3,\dots & 1^*, 2, \dots & 1,2,\dots \end{array}$$

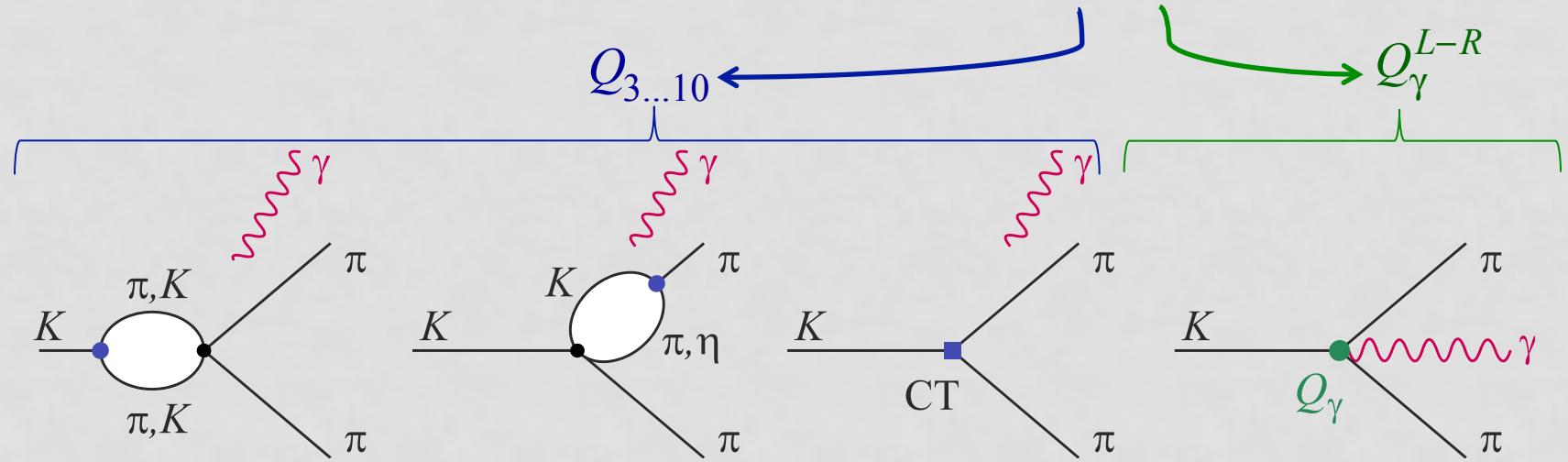
Dominantly CP-violating observables, with minimal **n** and **m** (largest rates):

Real photon(s):  $A_{CP}(K \rightarrow \gamma\gamma, K \rightarrow \pi\pi\gamma)$  (only magnetic)

Virtual photon:  $Br(K_L \rightarrow \pi^0\gamma^* [\rightarrow \ell^+\ell^-])$  (electric and magnetic)

B. Anatomy of  $K^\pm \rightarrow \pi^\pm \pi^0 \gamma$ 

Meson processes lack the chirality flip → Automatic LD-SD factorization



- Good control over meson contributions (loops usually finite, no/small CTs)
- Non-local CP-violating effects ( $Q_{3 \rightarrow 10}^{u,d,s}$ ) → To be estimated using  $\varepsilon'/\varepsilon$
- Local CP-violating effects ( $Q_{\gamma}^{L\pm R}$ ) → Local chiral realization of the tensor currents

We got a probe for NP in the magnetic photon operator  $Q_{\gamma}^{L-R}$

## B. Short reminder

Isospin decomposition:

$$\mathcal{M}(K_1 \rightarrow \pi^+ \pi^-) = \sqrt{2} A_0 + A_2$$

$$\mathcal{M}(K_1 \rightarrow \pi^0 \pi^0) = \sqrt{2} A_0 - 2 A_2$$

$$\mathcal{M}(K^+ \rightarrow \pi^+ \pi^0) = 3/2 A_2$$

$\Delta I = 1/2$  rule:

$$\frac{\text{Re } A_0}{\text{Re } A_2} = 22.4 \equiv \omega^{-1}$$

CP-violation:  $\frac{\mathcal{M}(K_L \rightarrow \pi^+ \pi^-)}{\mathcal{M}(K_S \rightarrow \pi^+ \pi^-)} = \varepsilon + \varepsilon'$ ,  $\frac{\mathcal{M}(K_L \rightarrow \pi^0 \pi^0)}{\mathcal{M}(K_S \rightarrow \pi^0 \pi^0)} = \varepsilon - 2\varepsilon'$

$$\varepsilon = \varepsilon_{box} + i \frac{\text{Im } A_0}{\text{Re } A_0} \sim 10^{-3}, \quad \varepsilon' = i \frac{e^{i(\delta_2^0 - \delta_0^0)}}{\sqrt{2}} \omega \left[ \frac{\text{Im } A_2}{\text{Re } A_2} - \frac{\text{Im } A_0}{\text{Re } A_0} \right] \sim 10^{-6}$$

$\nearrow$

$\Delta S = 2$  mixing

$$\frac{\text{EW penguins } (Q_8)}{\text{QCD penguins } (Q_6)} = \frac{\text{Im } A_2}{\text{Im } A_0} \stackrel{SM?}{=} \Omega \omega = 0.35(15) \omega$$

## C. SD sensitivity from CP asymmetry

Ecker, Neufeld, Pich, '94

Direct charge asymmetry: interfering amplitudes + different weak &amp; strong phases.

$$\mathcal{M} = \varepsilon_{\mu}^*(k) \left[ E(p_{+,0}) \frac{p_0^{\mu} p_+ \cdot k - p_+^{\mu} p_0 \cdot k}{m_K^3} + M(p_{+,0}) \frac{i \varepsilon^{\mu\nu\rho\sigma} p_{+,0} p_{0,\rho} k_{\sigma}}{m_K^3} \right]$$

$$E(p_{+,0}) = E_{IB}(p_{+,0}) + E_{DE}(p_{+,0})$$

Low theorem:

$$E_{IB}(p_{+,0}) \sim \frac{\mathcal{M}(K^+ \rightarrow \pi^+ \pi^0)}{k \cdot p_+ k \cdot p_K}$$

Strong phase:  $I = 2, L = 0$ Weak phase:  $Q_{7 \rightarrow 10}$ 

Multipole expansion:

$$E_{DE}(p_{+,0}) = E_1(k) + E_2(k)(p_+ - p_0) \cdot k + \dots$$

Strong phase:  $I = 1, L = 1$ Weak phase:  $Q_{3 \rightarrow 10}, Q_{\gamma}^{L-R}$

### C. SD sensitivity from CP asymmetry

The CP-violating parameter is

$$\varepsilon'_{+0\gamma} = \frac{\text{Im } E_{DE}}{\text{Re } E_{DE}} - \frac{\text{Im } A_2}{\text{Re } A_2} \approx -\frac{2}{3} \frac{\sqrt{2} |\varepsilon'|}{\omega} \left[ 1 + \frac{\Omega}{1-\Omega} \omega \right] + 3 \text{Im } C_\gamma^{L-R}$$

MP,Smith, '11

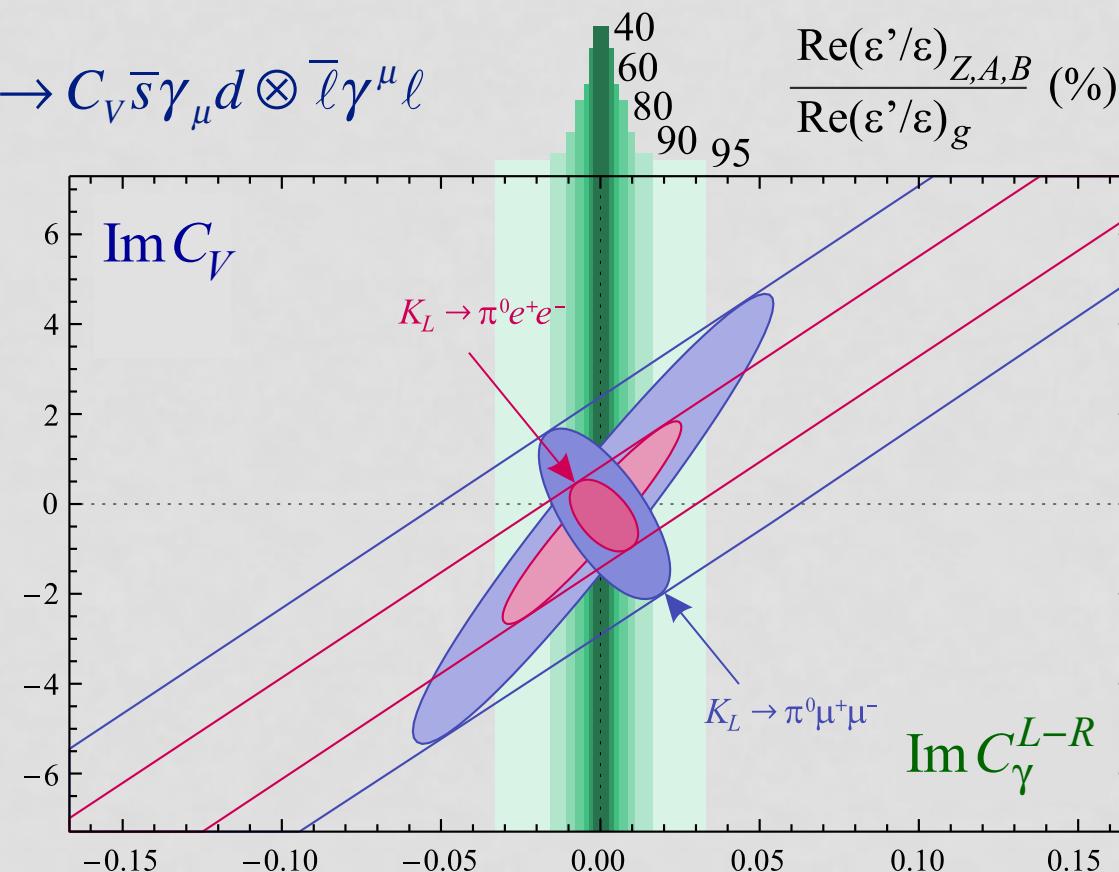
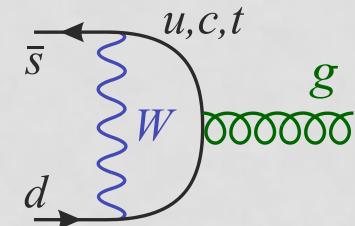
$$-0.6(3) \times 10^{-4} \quad +1.2(4) \times 10^{-4}$$

- The  $\Delta I = 1/2$  suppressed loops are enhanced by the  $\pi\pi$  contribution
- $\varepsilon'_{+0\gamma}$  is not  $\Delta I = 1/2$  suppressed, but its sensitivity to  $\Omega$  is !

Experimentally:  $\varepsilon'_{+0\gamma} = -0.21 \pm 0.34 \rightarrow \text{Im } C_\gamma^{L-R} \leq -0.08 \pm 0.13$   
 (NA48/2 '10)

Large room for NP effects !

## D. Correlation in the MSSM

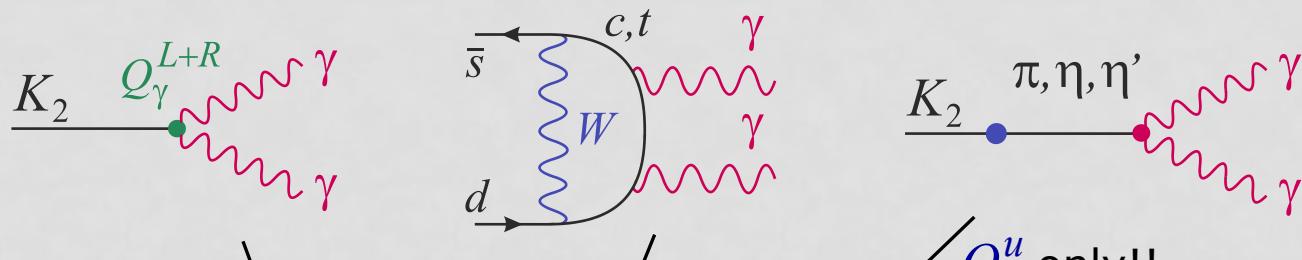
Buras, Colangelo, Isidori,  
Romanino, Silvestrini, '00Loop particles colored and charged  $\rightarrow$  Chromomagnetic penguinsLet us fix  $\text{Im } C_{\gamma}^{L+R} = \pm 3/2 \text{Im } C_g^{L-R}$  $Z, \gamma^* penguins \rightarrow C_V \bar{s} \gamma_{\mu} d \otimes \bar{\ell} \gamma^{\mu} \ell$ 

## The $K^0 \rightarrow \gamma\gamma$ decays

The CP-violating parameters are

$$\frac{\mathcal{M}(K_L \rightarrow \gamma\gamma_{||})}{\mathcal{M}(K_S \rightarrow \gamma\gamma_{||})} = \varepsilon + \varepsilon'_{||}, \quad \frac{\mathcal{M}(K_S \rightarrow \gamma\gamma_{\perp})}{\mathcal{M}(K_L \rightarrow \gamma\gamma_{\perp})} = \varepsilon + \varepsilon'_{\perp}$$

For the orthogonal polarization:



$$\varepsilon'_{\perp} = i \left[ \frac{\text{Im } C_{\gamma}^{L+R}}{2} + <10^{-7} + : 0 - \frac{\text{Im } A_0}{\text{Re } A_0} \right] \approx i \frac{\sqrt{2} |\varepsilon'|}{\omega(1-\Omega)} : 10^{-5} \quad 5 \times 10^{-5} \rightarrow 7 \times 10^{-4}$$

MP,Smith, '11

Gérard,Smith,Trine, '05

Directly measures the QCD penguins (hence  $\Omega$ ) and is  $\Delta I = 1/2$  enhanced.

1-  $K \rightarrow \pi v\bar{v}$  and  $K_L \rightarrow \pi^0 \ell^+ \ell^-$  decays are unique windows into the  $s \rightarrow d$  sector.

Essential to constrain (and discriminate among) New Physics models.

2- *Message to experimentalists*: Go for the  $K \rightarrow \pi v\bar{v}$  modes, they are the cleanest,

*But do not disregard other modes*:

$K_L \rightarrow \pi^0 \ell^+ \ell^-$  : Sensitive to a larger class of NP effects

$K_S \rightarrow \pi^0 \ell^+ \ell^-$ ,  $K_L \rightarrow \pi^0 \gamma\gamma$ ,  $K_{\ell 3}$  : Theoretical control for rare K decays

$K_S \rightarrow \pi^0 \gamma\gamma$ ,  $K^+ \rightarrow \pi^+ \gamma\gamma$  : Theoretical control over  $K_L \rightarrow \mu^+ \mu^-$

*Radiative decays could hold the key to (finally) control and probe  $\epsilon'/\epsilon$*

- Magnetic contribution under control thanks to  $K^+ \rightarrow \pi^+ \pi^0 \gamma$ .

- The  $K_{L,S} \rightarrow \gamma\gamma_\perp$  asymmetry could directly measure QCD penguins.